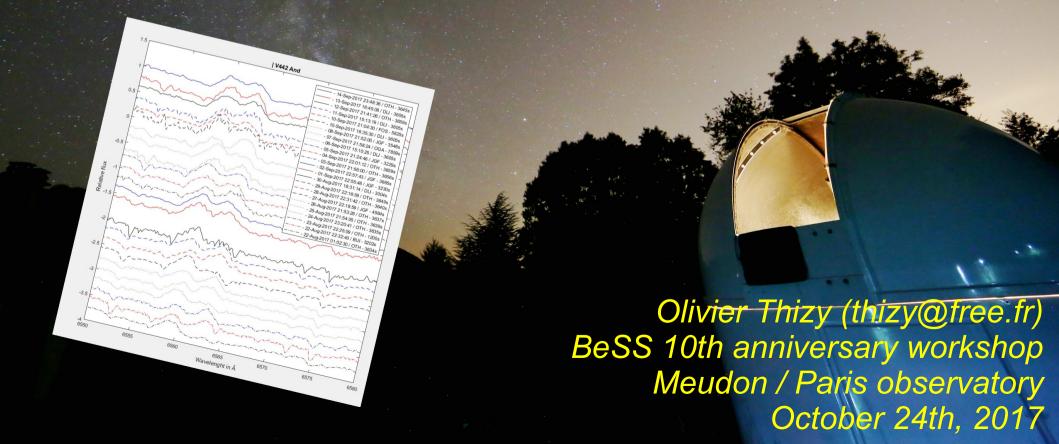
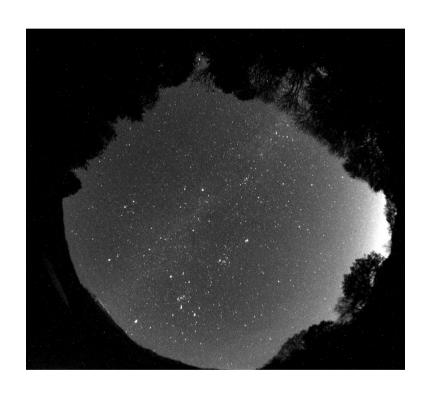
# V442 Andromedae August 2017 outburst



#### Content

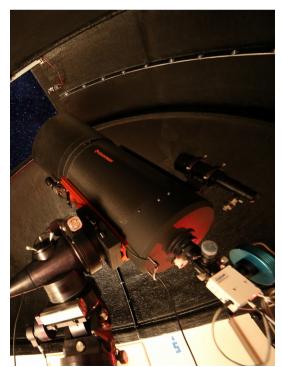


- Observatory of Belle Etoile
  - Overview
  - Be stars program
- V442 And ouburst
  - Discovery
  - Evolution
  - Possible scenario?
- Learning from this event

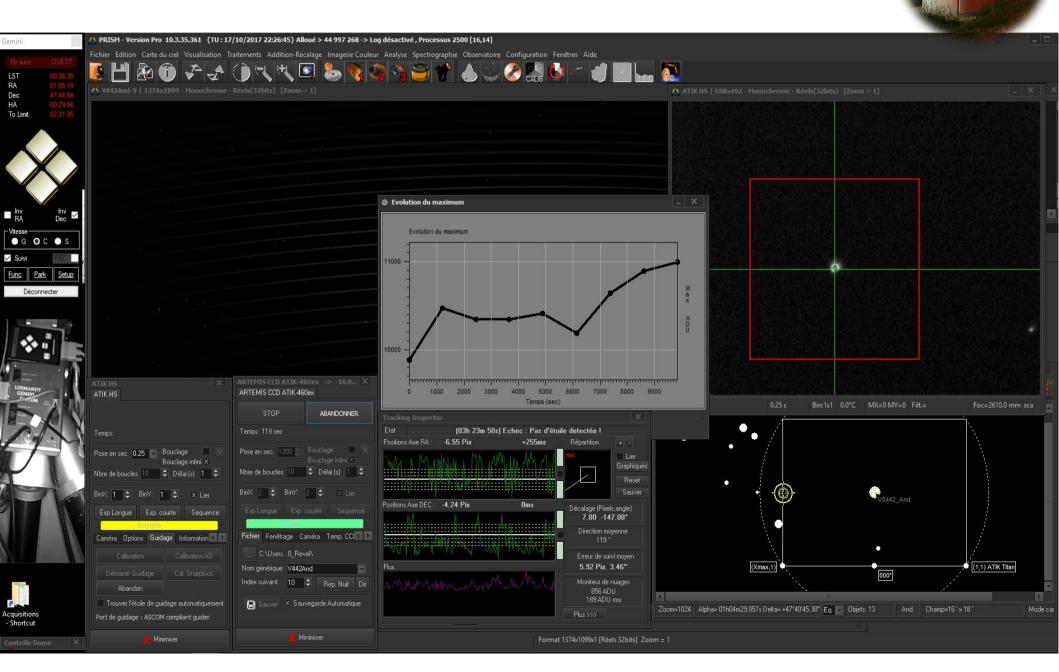
### Observatory of Belle Etoile

- In my backyard, 840m elevation
  - Average sky, poor seeing
- 2.7m dome, 0.28m telescope
- R~10000 echelle spectrograph
- All equipment operated with PRISM v10 software
- Since July 2015 "opening"
  - 100 nights, 725 hours observations
  - 427h total exposure time (59% "efficiency")





# Typical observing screen



## "BeSS" stars program



- Around 660 spectra (14000 orders) in BeSS
- 400+ spectra (~9500 orders) since July 2015
  - 100+ spectra of V442 And
  - Note: 74% of my spectra are "BeSS" stars
- ARASBeAm web / Robot to select targets
  - Look usually for Mag < 7-8 & DEC > -10°
  - Max individual exposure time 20min
  - Limit total exp. time 2h or **SNR** ~100 near H $\alpha$
- Manual data reduction with ISIS 5.8.0 (C.Buil)
  & BeSS upload during the night

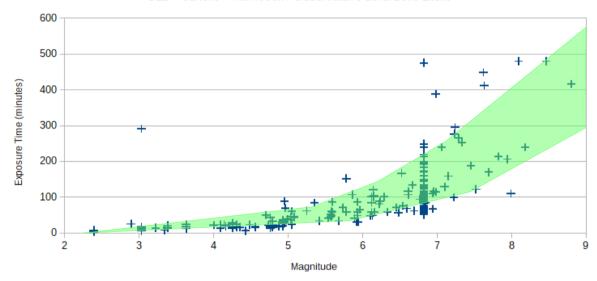
## Be stars spectra performances





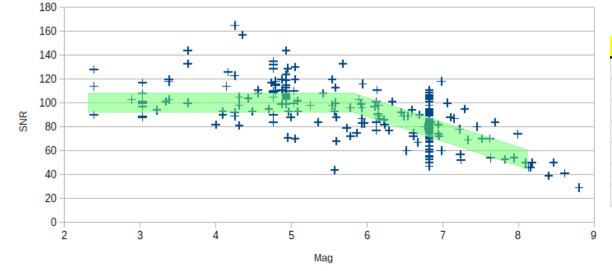
Be stars - Exposure Time (min) to reach SNR=100 Vs Magnitude

C11 + echelle + Atik460ex / Observatoire de la Belle Etoile



Be stars - SNR obtained per V Mag

C11 + echelle + Atik460ex / Observatoire de la Belle Etoile



| Date     | Etoile    | Exp.(s) | NbExp. | Max_kADU | SNR2 | Mag(Be) |
|----------|-----------|---------|--------|----------|------|---------|
| 19/10/17 | 28 Cyg    | 800     | 3      | 45       | 110  | 4,93    |
| 19/10/17 | V2163 Cyg | 1200    | 3      | 11       | 74   | 6,94    |
| 19/10/17 | V442 And  | 1200    | 5      | 11       | 82   | 6,82    |
| 19/10/17 | V442 And  | 1200    | 6      | 11       | 84   | 6,82    |
| 19/10/17 | V413 Aur  | 1200    | 8      | 4        | 46   | 8,13    |
| 19/10/17 | V715 Mon  | 1200    | 4      | 12       | 89   | 6,15    |
| 19/10/17 | OT Gem    | 1200    | 3      | 13       | 92   | 6,45    |

# Current / Futur Improvements



Bad Weather

- Remote observation since sep. 16th, 2017
- AllSky camera added last week
- Next steps: more scripting & automation
- Improve "efficiency": 78% in sep.-oct. 2017
- Automate Reduction / Archival / Dashboard ?

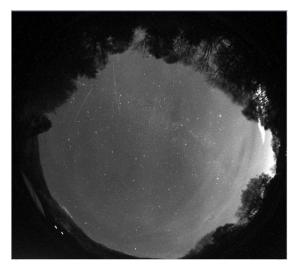
Non available

> Non 'BeSS'

**OVH** 

Photon time







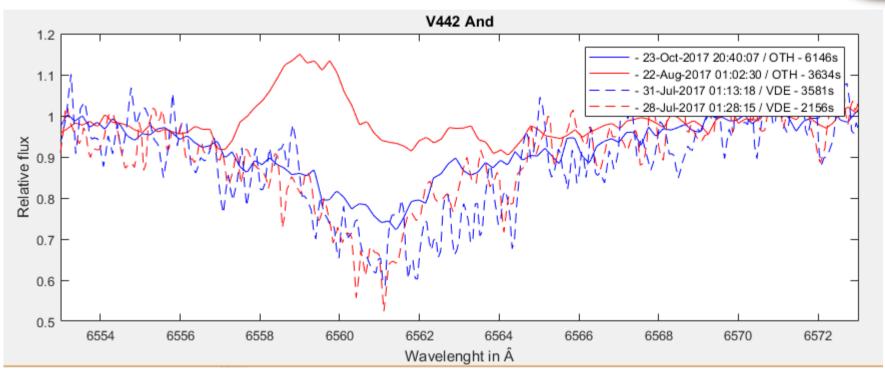
## V442 And (HD6226)



- First emission detected by McCollum et al. (2000)
- Bozic & Harmanec (2004) article on V442 And:
  - Star rotation 2.61507 days
  - 24-29 days period suspected
  - 630 days irregular period between outburst?
  - Teff 17000K; 5 solar masses; 11 solar radius
  - Inclination of 19° (almost pole on)

### V442 And outburst discovery



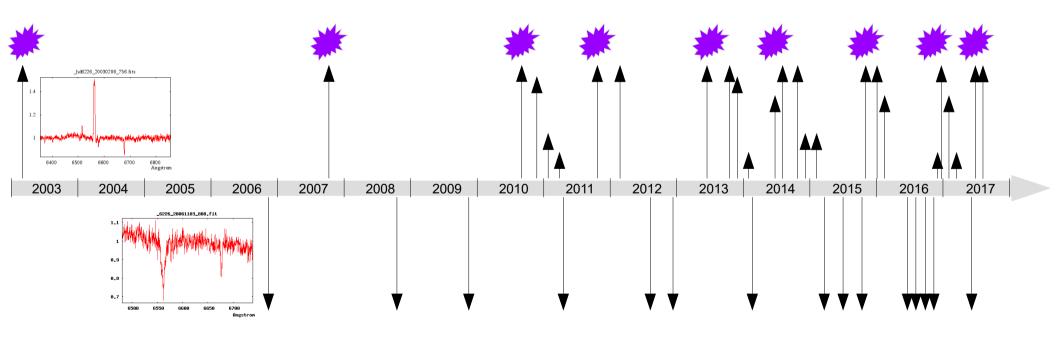


- Alert through Spectro-L Yahoo discussion group
- Regular status/infos on ARAS forum
- ARASBeAm "ideal period" quickly changed to 1d

## BeSS history

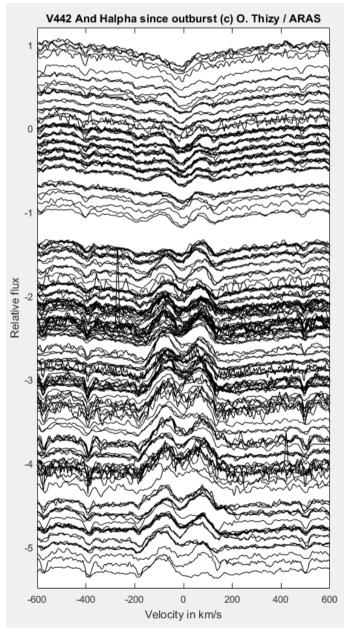


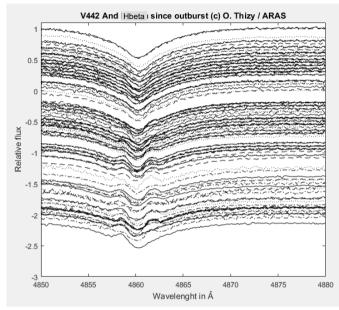
- Emission phases seem to last 3 to 6 months
- Intervals (quiescence): 8 to ~21 months



# 200+ spectra taken





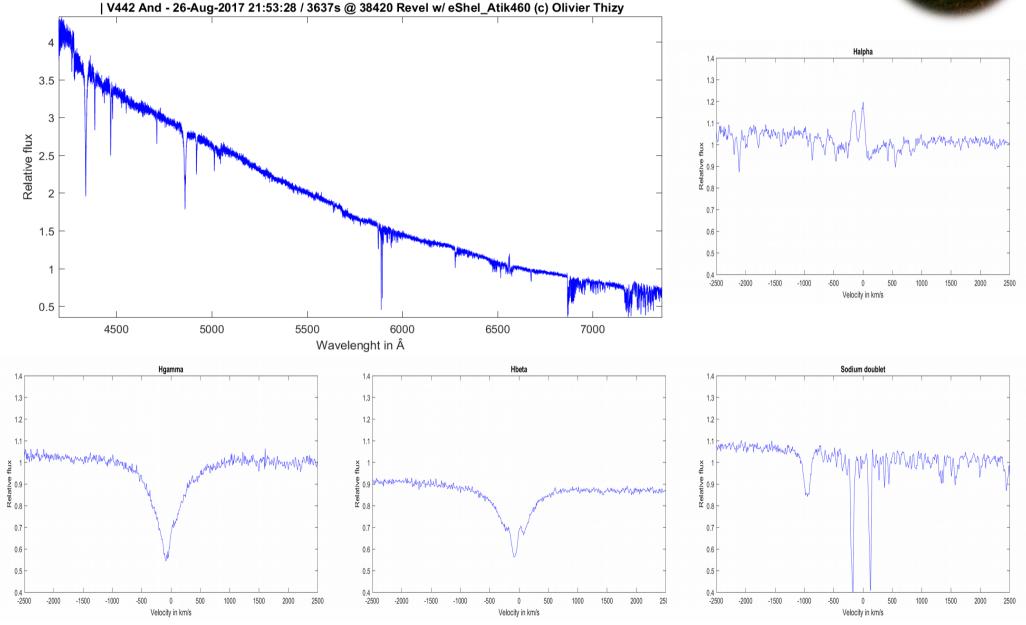


|     | Total expos       | <b>24</b> 5 | hours |     |           |
|-----|-------------------|-------------|-------|-----|-----------|
|     |                   |             |       |     |           |
|     |                   |             |       |     |           |
|     | Observer          | Spectro.    | #     | SNR | AvgExp(s) |
| OTH | Olivier Thizy     | echelle     | 110   | 85  | 4124      |
| DLI | Dong Li           | Littrow     | 44    | 40  | 4032      |
| JGF | Juan Guarro Flo   | echelle     | 32    | 106 | 3401      |
| OGA | Olivier Garde     | echelle     | 8     | 113 | 7557      |
| FOS | Patrick Fosanelli | Littrow     | 6     | 100 | 6242      |
| BUI | Christian Buil    | echelle     | 4     | 41  | 4089      |
| JNT | Jean-Noël Terry   | Littrow     | 3     | 57  | 3191      |
| VDE | Valérie Desnoux   | Littrow     | 2     | 50  | 3003      |
| ADB | Arnold de Bruin   | Littrow     | 1     | 165 | 5405      |
|     | Total Nb of sp    | 210         |       |     |           |

| ID   | File                            | Observer | Exp(s) | SNR | JD                   | V    | Vo      | R    | Ro             | Hel6678 |
|------|---------------------------------|----------|--------|-----|----------------------|------|---------|------|----------------|---------|
| 121  | _v442and_20170924_962_full      | OTH      | 3671   | 83  | 2458021,48295        | 0,26 | 6559,96 | 0,21 | 6562,48        | 6676,74 |
| 122  | _v442and_20170924_948_full_BUI  | BUI      | 7400   | 117 | 2458021,49094        | 0,28 | 6559,93 | 0,20 | 6562,50        | 6676,60 |
| 123  | _v442and_20170925_005_full      | OTH      | 2436   | 83  | 2458021,51870        | 0,27 | 6559,87 | 0,21 | 6562,49        | 6676,64 |
| 124  | _v442and_20170925_582_DongLi    | DLI      | 4206   | 21  | 2458022,10670        | 0,25 | 6559,45 | 0,19 | 6562,49        | 0,00    |
| 125  | _v442and_20170925_955_full      | OTH      | 7302   | 70  | 2458022,49708        | 0,26 | 6559,57 | 0,20 | 6562,18        | 6676,48 |
| 126  | _v442and_20170926_836_full      | OTH      | 1200   | 58  | 2458023,34263        | 0,27 | 6559,42 | 0,25 | 6562,67        | 6676,91 |
| 127  | _v442and_20170926_894_JGF       | JGF      | 4566   | 133 | 2458023,41998        | 0,27 | 6559,45 | 0,24 | 6562,77        | 6676,87 |
| 128  | _v442and_20170926_904_full_OGA  | OGA      | 7254   | 106 | 2458023,44631        | 0,28 | 6559,42 | 0,24 | 6562,89        | 6677,07 |
| 129  | _v442and_20170927_890_full      | OTH      | 6145   | 67  | 2458024,42516        | 0,27 | 6559,75 | 0,22 | 6563,09        | 6676,47 |
| 130  | _v442and_20170927_961_full      | OTH      | 6158   | 79  | 2458024,49678        | 0,26 | 6559,75 | 0,22 | 6562,88        | 6676,43 |
| 131  | _v442and_20170928_841_full      | OTH      | 3648   | 76  | 2458025,36252        | 0,19 | 6559,73 | 0,26 | 6562,53        | 6676,55 |
| 132  | _v442and_20170928_871_J. Guarro | JGF      | 3021   | 165 | 2458025,38804        | 0,20 | 6559,76 | 0,26 | 6562,50        | 6676,57 |
| 133  | _v442and_20170928_884_full      | OTH      | 3669   | 100 | 2458025,40527        | 0,18 | 6559,76 | 0,26 | 6562,58        | 6676,62 |
| 134  | _v442and_20170928_927_full      | OTH      | 3671   | 84  | 2458025,44817        | 0,18 | 6559,72 | 0,25 | 6562,57        | 6676,66 |
| 135  | _v442and_20170928_970_full      | OTH      | 4907   | 64  | 2458025,49832        | 0,19 | 6559,54 | 0,26 | 6562,60        | 6676,69 |
| 136  | _v442and_20170929_827_full      | OTH      | 3651   | 87  | 2458026,34792        | 0,15 | 6559,85 | 0,23 | 6562,82        | 6676,92 |
| 137  | v442and_20170929_869_full       | OTH      | 3647   | 71  | 2458026,39042        | 0,15 | 6559,80 | 0,24 | 6562,82        | 6676,92 |
| 138  | v442and_20170929_897_JGF        | JGF      | 4235   | 133 | 2458026,42170        | 0,16 | 6559,48 | 0,23 | 6562,91        | 6676,82 |
| 139  | v442and_20170929_912_full       | OTH      | 3680   | 73  | 2458026,43329        | 0,15 | 6559,62 | 0,24 | 6562,84        | 6676,88 |
| 140  | v442and_20170929_899_full       | OGA      | 7253   | 123 | 2458026,44056        | 0,15 | 6559,68 | 0,24 | 6563,00        | 6676,97 |
| 141  | v442and_20171003_779_full_OGA   | OGA      | 7257   | 117 | 2458030,32098        | 0,19 | 6559,24 | 0,17 | 6562,87        | 6676,41 |
| 142  | v442and_20171003_919_J. Guarro  | JGF      | 3621   | 80  | 2458030,43994        | 0,19 | 6559,23 | 0,17 | 6562,95        | 6676,49 |
| 143  | v442and_20171004_834_full       | OTH      | 14842  | 75  | 2458031,42015        | 0,15 | 6559,51 | 0,17 | 6563,31        | 6677,05 |
| 1/// | w442and 20171006 818 full       | OTH      | £110   | 102 | <b>24E8U32 3E328</b> | 0.14 | CEEO EU | በ 10 | <b>6560 77</b> | EE7E 12 |

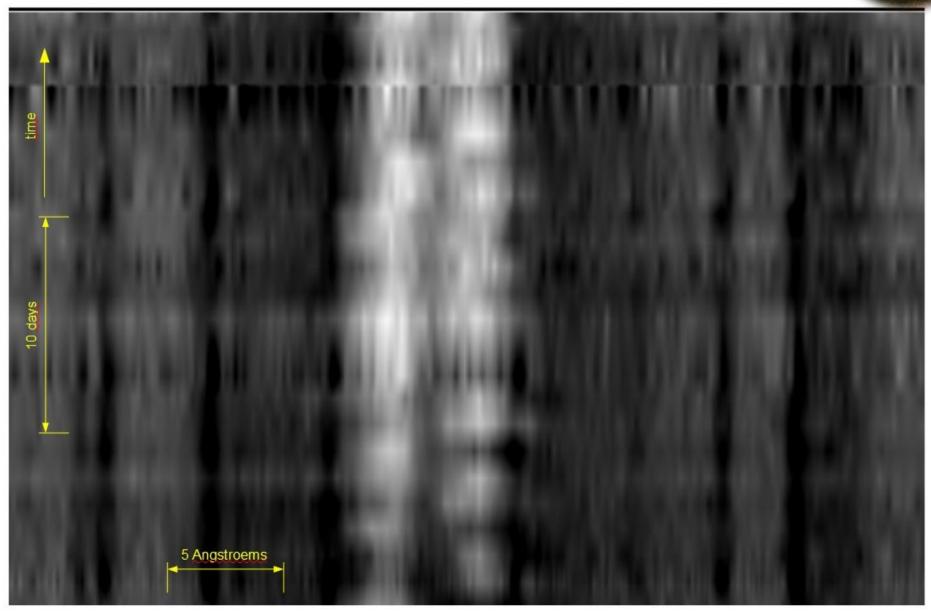
# Exemple of an echelle spectrum





# $H\alpha$ evolution — periodic ?

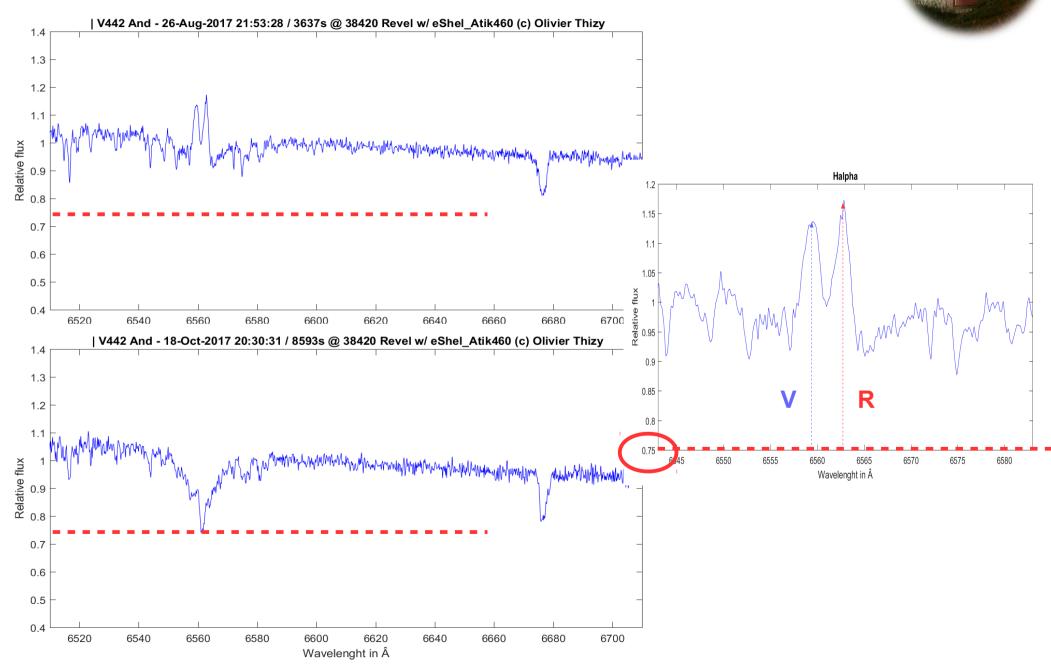




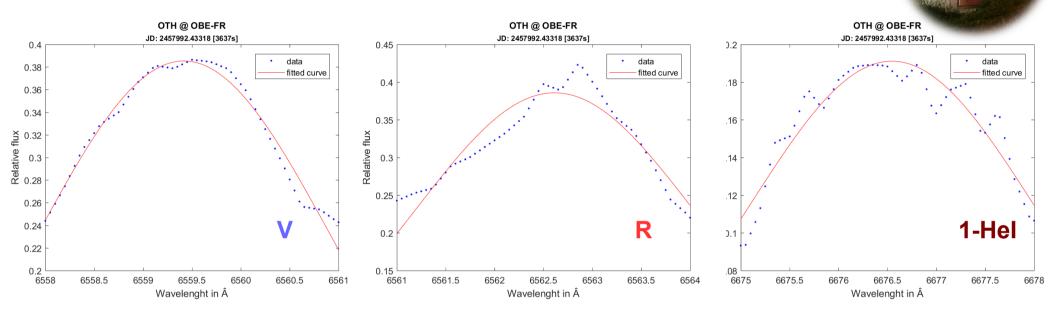
Based on ARAS/BeSS spectra: O. Thizy, J.Guarro Flo, O. Garde, Dong Li, C. Buil, V. Desnoux, P. Fosanelli, J.N. Terry

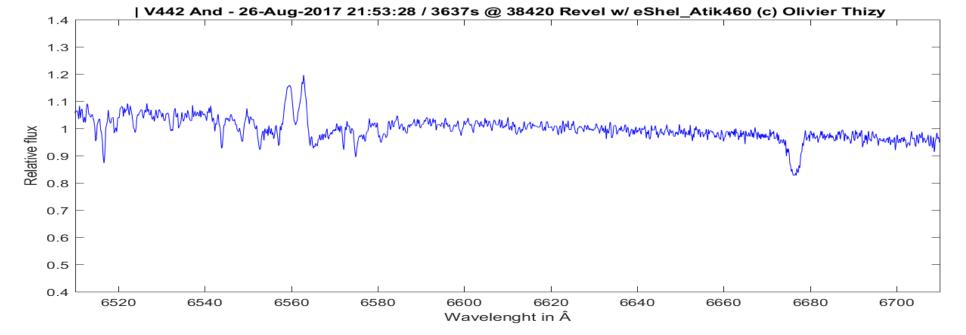
## Measuring V & R peaks





# MatLab gaussian fits



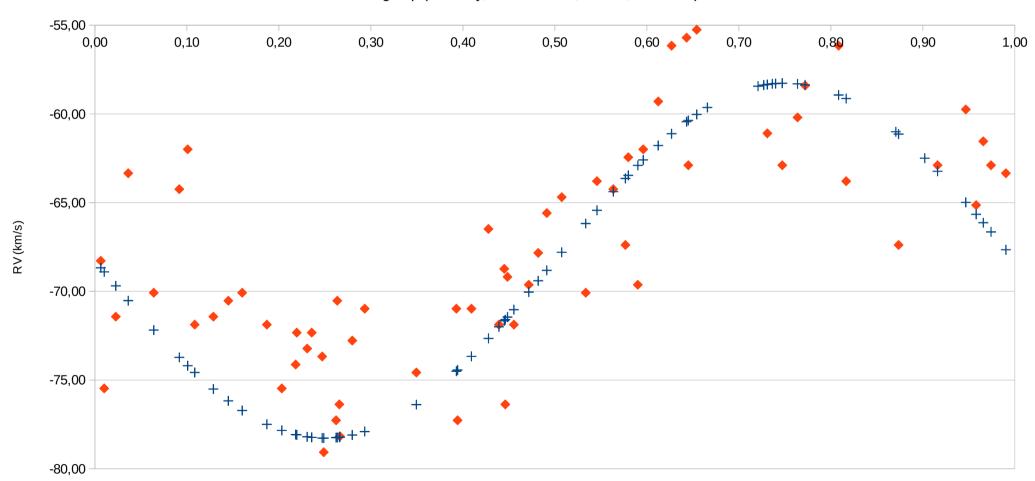


## He I: stellar rotation period



He I 6678 Radial Velocity

ARAS group (O. Thizy, J. Guarro Flo, C. Buil, O. Garde)

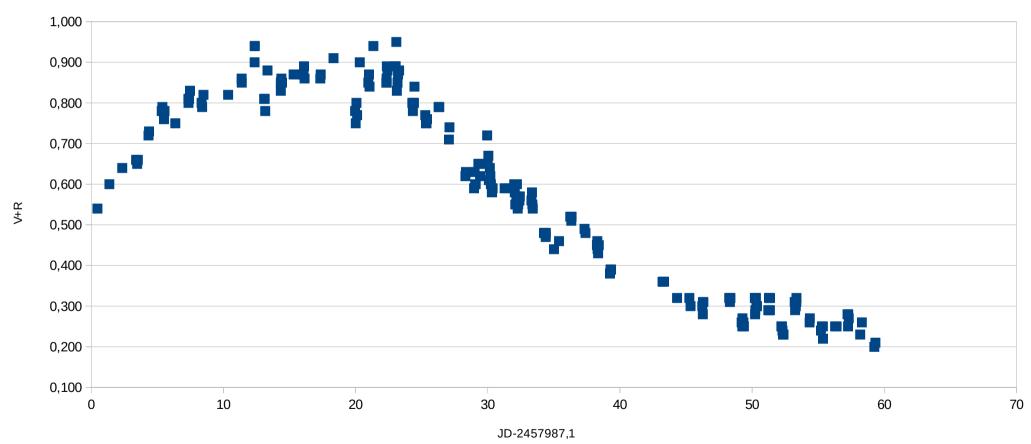


Phase (T0: 2457986,9 / P=2,6217 days)

### V+R evolution



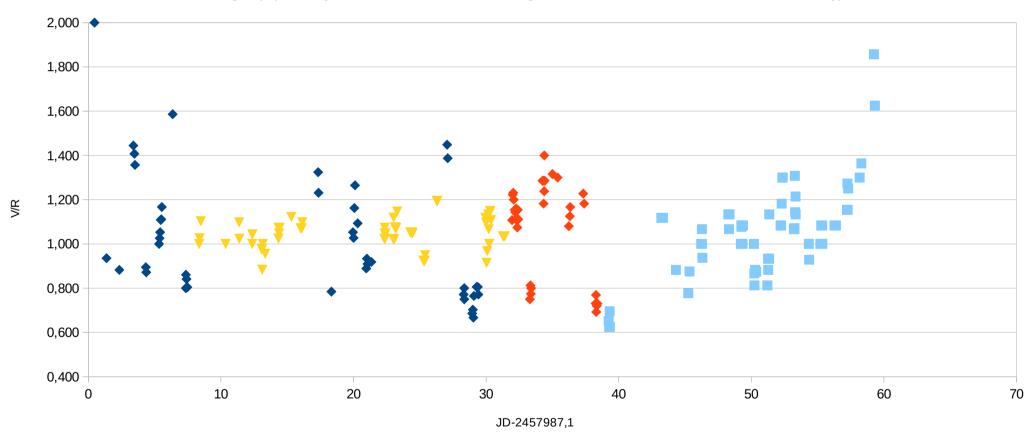
V442 And - V+R evolution after outburst discovery



### V/R evolution



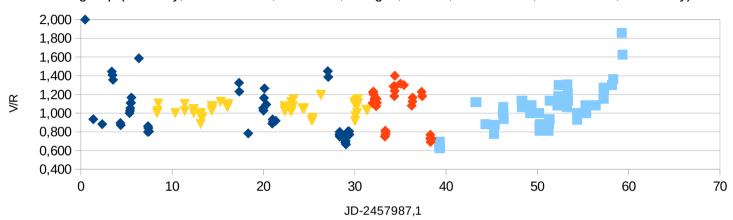
V442 And - V/R evolution after outburst discovery



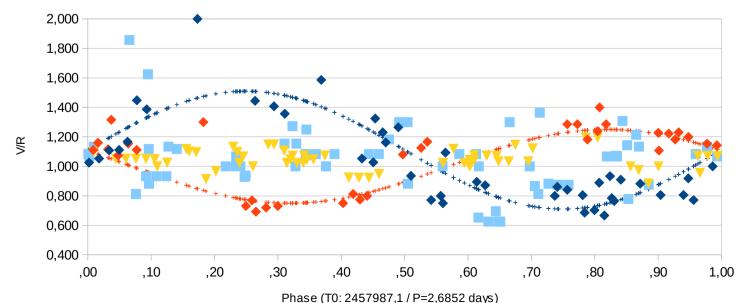
### V/R "phased" evolution

V442 And - V/R evolution after outburst discovery

ARAS group (O. Thizy, J. Guarro Flo, O. Garde, Dong Li, C. Buil, P. Fosanelli, V. Desnoux, J.N. Terry)



V42 And - V/R evolution after outburst discovery



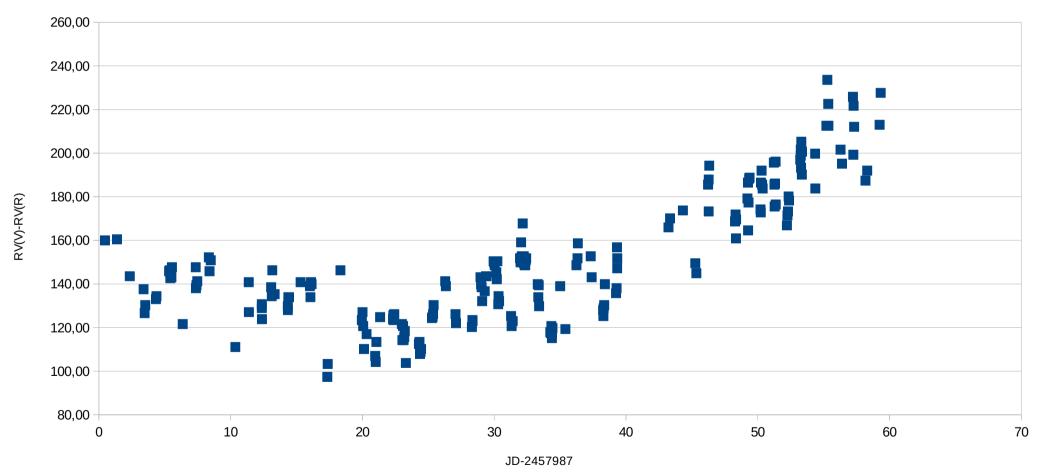


- A blob is ejected over the photosphere, one on each side (august 21st spectrum)
  - Material remains close to photosphere ~8d
  - System remains stable for about 5 days
- Another ejection ~4d + 4d
- Another ejection ~3d + 2d
- Then another ejection but with another phase
- Emission is now fainter and V/R more erratic

## V & R peaks separation



V442 And - V-R Radial Velocities after outburst discovery



## Learning from this event



- Checking spectra quickly & community alert is important
- Good (but not excellent) response from the observing community: Europe, China
- Lack of observers in USA/America's time zone
- Frequent update on ARAS forum and Yahoo Spectro-L discussion group helps for the group motivation
- Good reaction from Coralie Neiner & Noël Richardson: motivation, suggestions, explanations... THANKS!
- A publication is beeing prepared with additional observations – to be continued...

## Quotes from Noël Richardson



- Early on: "The quick V/R changes seem consistent with a blob ejected from the star and in orbit just above the stellar surface. This should make a very interesting project. Many thanks to you for arranging the spectroscopy and to all of the observers for their time, dedication, and enthusiasm."
- 100 spectra: "this seems like a rare opportunity to trace the mechanisms of how small ejections in Be stars operate to build a disk. Please keep observing!"
- 200 spectra: "From what I've seen, we haven't had such a great set of data on this (or any other) Be star on the onset and disappearance of a disk before."

